

Long term measurements of the plasmashet composition: implications on ionospheric outflow

Romain Maggiolo (1), Maria Hamrin (2), Gael Cessateur (1), Johan De Keyser (1), Herbert Gunell (1,2), Lukas Maes (3), Timo Pitkänen (4,2)

(1) BIRA/IASB, Brussels, Belgium (romain.maggiolo@aeronomie.be), (2) Umeå University, Umeå, Sweden, (3) Max Planck Institute for Solar System Research, Göttingen, Germany, (4) Institute of Space Science, Shandong University, Weihai, China

The plasmashet consists of a mixture of ions originating from the solar wind and from the ionosphere. Contrary to the solar wind, ionospheric ions contain a significant amount of O^+ ions which can be used as tracers of the ionospheric material. The ionospheric ion outflow rate, its composition and the transport of ionospheric ions from the ionosphere to the plasmashet are modulated by the solar wind conditions. This modulation of the ionospheric source impacts the plasmashet composition, in particular the amount of O^+ ions, with some time delay related to ionospheric ion outflow and transport processes.

Using long-term O^+ and H^+ density measurements from the CODIF experiment onboard the Cluster spacecraft we investigate the spatial distribution of O^+ and H^+ ions in the Earth plasmashet and its dependence on solar wind and geomagnetic activity. This allows the identification of the solar wind parameters that have the strongest impact on the plasmashet O^+ density - and thus on ionospheric ion outflow and transport - as well as the associated response times. We discuss the implication of these results for our understanding of ionospheric ion sources and outflow processes.