

Why an intrinsic magnetic field does not protect a planet against atmospheric escape

Herbert Gunell,^{1,2} Romain Maggiolo,¹ Hans Nilsson,³ Gabriella Stenberg Wieser,³
Rikard Slapak,⁴ Jesper Lindkvist,² Maria Hamrin,² and Johan De Keyser,¹

¹Royal Belgian Institute for Space Aeronomy, Brussels, Belgium

²Umeå University, Umeå, Sweden

³Swedish Institute of Space Physics, Kiruna, Sweden

⁴EISCAT Scientific Association, Kiruna, Sweden

The presence or absence of a magnetic field determines the nature of how a planet interacts with the solar wind and what paths are available for atmospheric escape.¹ Magnetospheres form both around magnetised planets, such as Earth, and unmagnetised planets, like Mars and Venus,² but it has been suggested that magnetised planets are better protected against atmospheric loss.³ However, the observed mass escape rates from these three planets are similar,^{4,5} putting the hypothesis into question. Modelling the effects a planetary magnetic field has on the major atmospheric escape processes, we show that the escape rate can be higher for magnetised planets over a wide range of magnetisations due to escape of ions through the polar caps and cusps.⁶ Thus, contrary to what previously has been believed,^{1,3,7} magnetisation is not a sufficient condition for protecting a planet from atmospheric loss. Estimates of the atmospheric escape rates from exoplanets⁷ must therefore address all escape processes and their dependence on the planet's magnetisation.

¹R. Lundin, H. Lammer, and I. Ribas, “Planetary magnetic fields and solar forcing: Implications for atmospheric evolution,” *Space Science Reviews*, vol. 129, pp. 245–278, Mar. 2007.

²C. T. Russell, “Planetary magnetospheres,” *Rep. Prog. Phys.*, vol. 56, pp. 687–732, 1993.

³K. Seki, R. C. Elphic, M. Hirahara, T. Terasawa, and T. Mukai, “On atmospheric loss of oxygen ions from earth through magnetospheric processes,” *Science*, vol. 291, pp. 1939–1941, Mar. 2001.

⁴H. Lammer, *Origin and Evolution of Planetary Atmospheres: Implications for Habitability*. Springer Briefs in Astronomy, Berlin: Springer, 2013.

⁵M. André and C. M. Cully, “Low-energy ions: A previously hidden solar system particle population,” *Geophys. Res. Lett.*, vol. 39, p. 3101, Feb. 2012.

⁶H. Gunell, R. Maggiolo, H. Nilsson, G. Stenberg Wieser, R. Slapak, J. Lindkvist, M. Hamrin, and J. De Keyser, “Why an intrinsic magnetic field does not protect a planet against atmospheric escape,” *Astronomy & Astrophysics*, 2018. submitted.

⁷H. Lammer, H. I. M. Lichtenegger, Y. N. Kulikov, J.-M. Grießmeier, N. Terada, N. V. Erkaev, H. K. Biernat, M. L. Khodachenko, I. Ribas, T. Penz, and F. Selsis, “Coronal Mass Ejection (CME) Activity of Low Mass M Stars as An Important Factor for The Habitability of Terrestrial Exoplanets. II. CME-Induced Ion Pick Up of Earth-like Exoplanets in Close-In Habitable Zones,” *Astrobiology*, vol. 7, pp. 185–207, Feb. 2007.